

30 years of many-body in Spain: looking into the future

of mathematical methods in many-body physics

(Friday, September 24th, 2004, 18.45 hours)

Spiked A-body anharmonic oscillators in a strong-repulsion perturbative picture

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I. General subject: A -particle bound-state problems

$$\hat{H}\Psi(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_A) = E\Psi(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_A).$$

Here, their first nontrivial case with $A = 3$ and $x_j \in I\!\!R$,

$$\left[-\sum_{i=1}^3 \frac{\partial^2}{\partial x_i^2} + \hat{W} \right] \Psi = E\Psi,$$

$$\hat{W} = V(x_1 - x_2) + V(x_2 - x_3) + V(x_3 - x_1).$$

The best known Calogero's solvable option

$$V^{(spiked)}(x_i - x_j) = \frac{\nu(\nu + 1)}{(x_i - x_j)^2} + \omega^2 (x_i - x_j)^2$$

Here, its $L, K = 1, 2, \dots$ generalizations

$$V(x_i - x_j) = \alpha^2 (x_i - x_j)^L + \beta^2 (x_i - x_j)^{-K}$$

and their superpositions.

II. The method

Taylor identities in the strong-repulsion regime:

(1) SHO has a minimum at $R = [\nu(\nu + 1)/\omega^2]^{1/4}$,

$$\begin{aligned}
 V_{eff}(r) &= \frac{\nu(\nu + 1)}{r^2} + \omega^2 r^2 \\
 &= V_{eff}(R) + \frac{1}{2}(r - R)^2 V''_{eff}(R) + \mathcal{O}[(r - R)^3] \\
 &= 2\omega^2 R^2 + 4\omega^2(r - R)^2 + \mathcal{O}\left[\frac{(r - R)^3}{R}\right], \\
 &= \text{solv. HO well at } 1/R \approx 1/\sqrt{\nu} \ll 1.
 \end{aligned}$$

(2) Taylor identities in the quartic case with $r = R + x$

clarify the determination of R as a root of $\mathcal{O}(x)$ term, etc.,

$$\begin{aligned}
 V_{eff}(r) = & \left[\frac{\nu (\nu + 1)}{R^2} + \omega^2 R^2 + \lambda R^4 \right] + \\
 & + 2 \left[\omega^2 R - \frac{\nu (\nu + 1)}{R^3} + 2 \lambda R^3 \right] x + \\
 & + \left[\omega^2 + 3 \frac{\nu (\nu + 1)}{R^4} + 6 \lambda R^2 \right] x^2 + \\
 & + 4 \left[-\frac{\nu (\nu + 1)}{R^5} + \lambda R \right] x^3 + O(x^4).
 \end{aligned}$$

(3) An unsolvable example of UE \longleftrightarrow RHO \longleftrightarrow SHO,

$$W(r) = F r^3 + \frac{G}{r^3}, \quad r, F, G > 0.$$

$$G \gg 1, r_{min} = R = (G/F)^{1/6} \gg 1$$

$$W(R + \xi) = 2 \sqrt{G F} + 9 \sqrt{G F} \frac{\xi^2}{R^2} - \dots .$$

$$W_0(R_0 + \xi) = F_0 (R_0 + \xi)^2 + \frac{G_0}{(R_0 + \xi)^2}$$

$$F_0 = \frac{9}{4} \sqrt[6]{G F^5}, \quad G_0 = \frac{4}{9} \sqrt[6]{G^5 F}.$$

III. The $A = 3$ models in Jacobi coordinates

$$\begin{pmatrix} Z \\ X \\ Y \end{pmatrix} = \begin{pmatrix} 1/\sqrt{3} & 1/\sqrt{3} & 1/\sqrt{3} \\ 1/\sqrt{2} & -1/\sqrt{2} & 0 \\ 1/\sqrt{6} & 1/\sqrt{6} & -2/\sqrt{6} \end{pmatrix} \begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix}$$

$$\begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix} = \begin{pmatrix} 1/\sqrt{3} & 1/\sqrt{2} & 1/\sqrt{6} \\ 1/\sqrt{3} & -1/\sqrt{2} & 1/\sqrt{6} \\ 1/\sqrt{3} & 0 & -2/\sqrt{6} \end{pmatrix} \begin{pmatrix} Z \\ X \\ Y \end{pmatrix}.$$

with interactions

$$\hat{W}(x_1, x_2, x_3) = \sum_{m=-K}^L F_m W^{(m)}(x_1, x_2, x_3),$$
$$F_L = \alpha^2 > 0, \quad F_{-K} = \beta^2 > 0,$$

$$W^{(m)} = (x_1 - x_2)^m + (x_2 - x_3)^m + (x_3 - x_1)^m$$

and Schrödinger equation with $U^{(m)}(X, Y) \equiv W^{(m)}(x_1, x_2, x_3)$,

$$\left\{ -\frac{\partial^2}{\partial X^2} - \frac{\partial^2}{\partial Y^2} + \sum_{m=-K}^L F_m U^{(m)} - E \right\} \Phi = 0$$

A polar re-parametrization of the coordinates,

$$X = X(\varrho, \varphi) = \varrho \sin \varphi$$

$$Y = Y(\varrho, \varphi) = \varrho \cos \varphi$$

gives Schrödinger equation with $\Omega^{(m)}(\varrho, \varphi) \equiv U^{(m)}(X, Y)$,

$$\left\{ -\frac{\partial^2}{\partial \varrho^2} - \frac{1}{\varrho^2} \frac{\partial^2}{\partial \varphi^2} + \sum_{m=-K}^L F_m \Omega^{(m)} - E \right\} \Phi = 0$$

Note the “forgotten” separability for the
1st illustrative oscillator: quartic,

$$U^{(2)}(X, Y) = 3X^2 + 3Y^2$$

$$\Omega^{(2)}(\varrho, \varphi) = U^{(2)}[X(\varrho, \varphi), Y(\varrho, \varphi)] = 3\varrho^2$$

$$U^{(4)}(X, Y) = \frac{9}{2} (X^2 + Y^2)^2 = \frac{9}{2} \varrho^4$$

A guiding return to Calogerian $U^{(SHO)}(X, Y)$ reveals that

(1) its $A = 3$ spiked part = a trigonometric identity:

$$\begin{aligned} \frac{1}{2X^2} + \frac{1}{2(X - \sqrt{3}Y)^2} + \frac{1}{2(X + \sqrt{3}Y)^2} &\equiv \\ \equiv \frac{1}{2\varrho^2} \left(\frac{1}{\sin^2 \varphi} + \frac{1}{\sin^2(\varphi + \frac{2}{3}\pi)} + \frac{1}{\sin^2(\varphi - \frac{2}{3}\pi)} \right) &\equiv \\ \equiv \frac{9}{2\varrho^2 \sin^2 3\varphi} \end{aligned}$$

and its Schrödinger equation = also separable.

(2) its angular Taylor is easy,

$$\frac{1}{\sin^2 3(\gamma + \pi/6)} = 1 + 9\gamma^2 + 54\gamma^4 + \frac{1377}{5}\gamma^6 + \dots$$

while the radial is no news, with $R = \sqrt[4]{\frac{3\nu(\nu+1)}{2\omega^2}}$ and

$$\begin{aligned} V_{eff}(\varrho) &= \frac{9\nu(\nu+1)}{2\varrho^2} + 3\omega^2\varrho^2 = \\ &= V_{eff}(R) + \frac{1}{2}(\varrho - R)^2 V''_{eff}(R) + \dots = \\ &= 6\omega^2 R^2 + 12\omega^2(\varrho - R)^2 + \dots \end{aligned}$$

(3) test passed - the approximants prove correct,

$$\begin{aligned}\Omega^{(SHO)}(R + \xi, \pi/6 + \eta/R) &\approx \\ &\approx 6\omega^2 R^2 + 12\omega^2 \xi^2 + 27\omega^2 \eta^2\end{aligned}$$

with $\mathcal{O}(\eta^m \times \xi^n) \approx 1/R^{m+n-2} \ll 1$ while

$$E - 6\omega^2 R^2 \approx$$

$$2\sqrt{3}\omega(2n+1) + 3\sqrt{3}\omega(2m+1) + \dots .$$

(4) Weyl chambers = cartesian, rotated $\pi/6$;

kinetic energy = *locally* cartesian =

$$= -\frac{\partial^2}{\partial \varrho^2} - \frac{1}{\varrho^2} \frac{\partial^2}{\partial \varphi^2} = \\ -\frac{\partial^2}{\partial \varrho^2} - \frac{1}{(R + \xi)^2} \frac{\partial^2}{\partial \varphi^2} = -\frac{\partial^2}{\partial \xi^2} - \left[1 + \mathcal{O}\left(\frac{\xi}{R}\right) \right] \frac{\partial^2}{\partial \eta^2}$$

to be verified for $V = \omega^2 (x_i - x_j)^2 + \gamma (x_i - x_j)^3 + \frac{\nu(\nu+1)}{(x_i - x_j)^2}.$

IV. Choose $L = 3$

1. fix $x_1 > x_2 > x_3$, i.e., $\varphi \in (0, \pi/3)$
2. deduce $\Omega^{(3)}(\varrho, \varphi) \geq 0$ for $\gamma > 0$
3. study the extreme $\omega = 0$
4. spectrum = discrete since two minima
at $\varphi_{\pm} = \pi/6 - \varphi_{\mp}$ in

$$\Omega^{(SC)}(\varrho, \varphi) = \alpha^2 \varrho^3 \sin 3\varphi + \frac{\beta^2}{\varrho^2 \sin^2 3\varphi}$$

with

$$\alpha^2 = \frac{3}{2}\gamma > 0, \quad \beta^2 = \frac{9}{2}\nu(\nu + 1) > 0$$

and with the ϱ -dependence rule

$$\partial_\varphi \Omega^{(SC)} = 0 \text{ and its form}$$

$$\sin^2 3\varphi_{\pm} = \frac{2\beta^2}{\alpha^2 \varrho^5}$$

implying an asymptotic growth,

$$\Omega^{(SC)} > \frac{3\varrho}{2} \left(2\alpha^4 \beta^2 \varrho \right)^{1/3}, \quad \varrho \geq \varrho_0 = \left(\frac{2\beta^2}{\alpha^2} \right)^{1/5}.$$

(5) general M in $\Omega^{(2M+1)}(\varrho, \varphi)$ at $\varrho \gg 1$:

$$\Omega = \alpha^2 \Omega^{(2M+1)} + \beta^2 \Omega^{(-2)} \sim \varrho^{4M/3},$$

is growing quickly at the small

$$\varphi_- = \pi/3 - \varphi_+ \sim \varrho^{-(2M+3)/3}.$$

(6) specific $K = 3$

two minima at $\varrho \gg 1$ and the three valleys merge at

$$\varrho = \varrho_0 = (2\beta^2/\alpha^2)^{1/5}$$

absolute minimum at $\varphi = \varphi_0 = \pi/6$ and

$$R = \varrho_0^{(min)} = \sqrt[5]{2\beta^2/(3\alpha^2)} \gg 1$$

Re-define $\beta^2 = 3\alpha^2 R^5/2$, $\varrho = R + \xi$,

$\varphi = \pi/6 + \eta/R$ and expand

$$\begin{aligned} \frac{1}{\alpha^2} \Omega^{(SCO)}(\varrho, \varphi) &= \varrho^3 \sin 3\varphi + \frac{\beta^2}{\alpha^2 \varrho^2 \sin^2 3\varphi} = \\ &= \frac{5}{2} R^3 + 9 R \eta^2 + \frac{15}{2} R \xi^2 - \\ &- \frac{81}{2} \xi \eta^2 - 5 \xi^3 + \frac{1}{R} \left(\frac{675}{8} \eta^4 + 27 \xi^2 \eta^2 + \frac{15}{2} \xi^4 \right) \\ &\quad + \mathcal{O}\left(\frac{1}{R^2}\right). \end{aligned}$$

Schrödinger equation

$$\left[-\frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} + \alpha^2 \left(\frac{15}{2} x^2 + 9 y^2 \right) + \mathcal{O}(\sigma^5) \right] \psi =$$

$$= \sigma^2 \left(E - \frac{5}{2} R^3 \right) \psi$$

with $\sigma = 1/\sqrt[4]{R}$, $\varrho = R + \sigma x$ and

$$\varphi = \pi/6 + \sigma y/R.$$

Leading-order energies: closed formula

$$E = E_{m,n} = \frac{5}{2} R^3 + \\ + \alpha \sqrt{\frac{15}{2}} R (2m+1) + 3\alpha \sqrt{R} (2n+1) + \mathcal{O}\left(\frac{1}{R^{3/4}}\right).$$

V. Potentials $U^{(m)}$ with even $m = 2M$

Complicated $U^{(6)}(X, Y) =$

$$\frac{33}{4} X^6 + \frac{45}{4} X^4 Y^2 + \frac{135}{4} X^2 Y^4 + \frac{27}{4} Y^6,$$

$$U^{(8)}(X, Y) = \frac{3}{8} (X^2 + Y^2) \times$$

$$\times (27 Y^6 + 225 X^2 Y^4 - 15 X^4 Y^2 + 43 X^6),$$

abbreviation

$$\Omega^{(2M)}(\varrho, \varphi) = U^{(2M)}[X(\varrho, \varphi), Y(\varrho, \varphi)]$$

compact formulae

$$\Omega^{(6)}(\varrho, \varphi) = \frac{3\varrho^6}{4} \left(9 + 2 \sin^2 3\varphi\right),$$

$$\Omega^{(8)}(\varrho, \varphi) = \frac{3\varrho^8}{8} \left(27 + 16 \sin^2 3\varphi\right),$$

$$\Omega^{(10)}(\varrho, \varphi) = \frac{27\varrho^{10}}{16} \left(9 + 10 \sin^2 3\varphi\right),$$

$$\Omega^{(12)}(\varrho, \varphi) =$$

$$\frac{729\varrho^{12}}{32} + \frac{81\varrho^{12}}{2} \sin^2 3\varphi + \frac{3\varrho^{12}}{4} \sin^4 3\varphi.$$

$$\text{In general, } \frac{\Omega^{(2M)}(\varrho, \varphi)}{\varrho^{2M}} =$$

$$\frac{3^M}{2^{M-1}} + c_1 \sin^2 3\varphi + \dots + c_N \sin^{2N} 3\varphi$$

where, formally,

$$N = \text{entier} \left[\frac{M}{3} \right], \quad M = 1, 2, \dots$$

and $\text{entier}[x]$ denotes the integer part of a real number x .

VI. Potentials $\Omega^{(m)}$ with odd $m = 2M + 1$

$$U^{(3)}(X, Y) = \frac{3}{2} \sqrt{2} X \left(X^2 - 3Y^2 \right),$$

$$U^{(5)}(X, Y) = \frac{15}{4} \sqrt{2} \left(X^2 - 3Y^2 \right) \left(X^2 + Y^2 \right) X,$$

$$U^{(7)}(X, Y) = \frac{63}{8} \sqrt{2} \left(X^2 - 3Y^2 \right) X \left(X^2 + Y^2 \right)^2,$$

$$U^{(9)}(X, Y) = X \left(X^2 - 3Y^2 \right) \times$$

$$\times \frac{3}{16} \sqrt{2} \left(81Y^6 + 279X^2Y^4 + 219X^4Y^2 + 85X^6 \right)$$

while, in polar coordinates,

$$\Omega^{(3)}(\varrho, \varphi) = \frac{3\varrho^3}{2} \sqrt{2} \sin 3\varphi,$$

$$\Omega^{(5)}(\varrho, \varphi) = \frac{15\varrho^5}{4} \sqrt{2} \sin 3\varphi,$$

$$\Omega^{(7)}(\varrho, \varphi) = \frac{63\varrho^7}{8} \sqrt{2} \sin 3\varphi,$$

$$\begin{aligned} \Omega^{(9)}(\varrho, \varphi) &= \frac{3\varrho^9}{16} \sqrt{2} \sin 3\varphi \left(81 + 4 \sin^2 3\varphi \right) = \\ &= \frac{3\varrho^9}{16} \sqrt{2} \sin 3\varphi (83 - 2 \cos 6\varphi), \end{aligned}$$

$$\Omega^{(11)}(\varrho, \varphi) = \frac{33}{32} \varrho^{11} \sqrt{2} \sin 3\varphi \left(27 + 4 \sin^2 3\varphi \right),$$

$$\Omega^{(13)}(\varrho, \varphi) = \frac{117}{64} \varrho^{13} \sqrt{2} \sin 3\varphi \left(27 + 8 \sin^2 3\varphi \right),$$

etc. Again, $m = 6N + 3$ or $m = 6N + 5$ or

$m = 6N + 7$ leads to the general rule

$$\begin{aligned} \frac{\Omega^{(2M+1)}(\varrho, \varphi)}{\varrho^{2M+1}} &= \\ &= (2M+1) \sqrt{2} \sin 3\varphi \times \\ &\times \left(\frac{3^{M-1}}{2^M} + d_1 \sin^2 3\varphi + \dots + d_N \sin^{2N} 3\varphi \right) \end{aligned}$$

where

$$N = \text{entier} \left[\frac{M-1}{3} \right], \quad M = 1, 2, \dots .$$

VII. Strongly singular potentials $\Omega^{(m)}$ with negative m

$$\Omega^{(-1)}(\varrho, \varphi) = -\frac{3}{\varrho \sqrt{2} \sin 3\varphi},$$

$$\Omega^{(-2)}(\varrho, \varphi) = \left(\frac{3}{\varrho \sqrt{2} \sin 3\varphi} \right)^2 = [\Omega^{(-1)}(\varrho, \varphi)]^2.$$

$$\Omega^{(-3)}(\varrho, \varphi) = [\Omega^{(-1)}(\varrho, \varphi)]^3 \left(1 - \frac{4}{9} \sin^2 3\varphi \right),$$

$$\Omega^{(-4)}(\varrho, \varphi) = [\Omega^{(-1)}(\varrho, \varphi)]^4 \left(1 - \frac{16}{27} \sin^2 3\varphi \right),$$

$$\Omega^{(-5)}(\varrho,\varphi)=\left[\Omega^{(-1)}(\varrho,\varphi)\right]^5\left(1-\frac{20}{27}\sin^23\varphi\right),$$

$$\begin{aligned} \Omega^{(-6)}(\varrho,\varphi) &= \left[\Omega^{(-1)}(\varrho,\varphi)\right]^6\times \\ &\times\left(1-\frac{904}{667}\sin^23\varphi+\frac{8248}{18009}\sin^43\varphi\right). \end{aligned}$$

VIII. SUMMARY:

(A) Why are the spiked oscillators interesting?

(1) the old Aguilera's and Guardiola's answer in JMP '91:

$$E_{(\text{spiked HO})}^{(\text{ground state})} \text{ at } K = 5/2$$

[= the E. M. Harrell's Ann. Phys. '77 $A = D = 1$ formula]

has the same structure as the many-bosonic energy

$$E = a + b\beta^2 + c\beta^4 \log\beta + \mathcal{O}(\beta^6)$$

(2) they offer insights via analytic continuation
within PT symmetric quantum mechanics:

(*) single-particle case: FGRZ, JPA '98

(*) three-particle case: ZT, JPA '02

(3) they may exhibit partial solvability, etc.

(4) they have immediate impact on perturbation methods:

BFZ PRA '89, AEG JMP '90, FGZ CJP '91

(*) facilitated Padé-type resummations: FGZ PRA '93

(*) feasible $\beta \gg 1$ expansions in the past (GSR NC '92, $A = 1$),
today ($A = 3$) and in the future ($A > 3$)

(B) What's shared by $A = 1, 2$ and $A = 3$?

(Σ 1) mathematics:

(*) a Lie-algebraic background (Calogero)

(Σ 2) computing methods:

(*) a rigorous treatment of divergent series

(*) feasibility of symbolic manipulations: ZYG JPA '03

(Σ 3) guidance towards $A > 3$

(*) expansions using artificial parameters: Z PRA '87 and

M. Znojil, *Low-lying spectra . . .*", JPA 36 (2004) 9929-41